## New deformed Heisenberg algebra from the $\mu\text{-}\mathrm{Deformation}$ based model of dark matter

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Recently, the  $\mu$ -Bose condensate (extending Bose-Einstein condensate (BEC) phase), occuring in certain  $\mu$ -deformed Bose gas model, was used as a model of dark matter [1]. It shows some virtues relative to the BEC DM model [2]: (i) prediction of the total mass of galactic DM halo is in a somewhat better agreement with observations than in the BEC DM model; (ii) the critical temperature of condensation  $T_c(\mu)$  is higher for  $\mu > 0$ , making the  $\mu$ -condensate phase more stable against possible heating sources.

For the study of galaxy rotation curves, a  $\mu$ -deformed Lane-Emden (LE) equation is used, and different versions of  $\mu$ -LE arise [3]. We study three versions of deformation, compare with observation, and find certain improvement. Two versions are mutually equivalent (admit same solution built from  $\mu$ -sine function). From the equivalence, we infer new  $\mu$ -deformed Heisenberg algebra for the radial position operator  $r \cdot$  and the radial  $\mu$ -derivative  $D_r^{(\mu)}$  as generators. There are two distinct forms of  $\mu$ -Heisenberg algebra: one involves usual commutator in the l.h.s. and nontrivial terms in the r.h.s.; the other has usual r.h.s.  $i\hbar$ , but operates with unusual  $\mu$ -deformation of the commutator in the l.h.s. of relation.

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